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Auxiliary Field Meson Model at Finite Temperature and DensityH. Kouno¹, T. Sakaguchi², K. Kashiwa², M. Hamada²,H. Tokudome¹, M. Matsuzaki^{2,3} and M. Yahiro²¹ *Department of Physics, Saga University, Saga 840-8502, Japan*² *Department of Physics, Kyushu University, Fukuoka 812-8581, Japan*³ *Department of Physics, Fukuoka University of Education,
Munakata, Fukuoka 811-4192, Japan***Abstract**

Starting from many quark interactions, we construct a nonlinear σ - ω model at finite temperature and density. The mesons are introduced as auxiliary fields. Effective quark-meson couplings are strongly related to effective meson masses, since they are derived simultaneously from the original many quark interactions. In this model, even if the effective ω -meson mass decreases due to the partial chiral restoration, the equation of state (EOS) of nuclear matter can become soft.

1 Introduction

The ω -meson is important for the nuclear structure. It is reported that the reduction of the effective ω -meson mass makes the nuclear matter EOS stiffer. [1] However, if we require that the ω -meson mean field be proportional to the baryon density, the effective ω -nucleon coupling also becomes smaller as the effective ω -meson mass becomes smaller and the EOS of nuclear matter becomes softer. [2]

In this paper, we show that, at finite temperature and density, effective meson-quark couplings are strongly related to effective meson masses, if the meson fields are introduced as auxiliary fields which consist of quarks and anti-quarks. Consequently, if the effective ω -meson mass decreases, the effective ω -quark (or nucleon) coupling decreases and the EOS of nuclear matter becomes softer. Therefore, even if the effective ω -meson mass decreases due to the partial chiral restoration, the EOS of nuclear matter can become soft in this model.

2 Auxiliary field method for nonlinear σ - ω model

In this section, using the auxiliary field method, [3,4] we construct a nonlinear σ - ω model. (For details, see the reference [5].) We start from the many quark interactions [4,5]

$$\begin{aligned}
\int dt V &= \sum_{m+n \geq 2} \frac{1}{m!n!} \int d^4x_1 \cdots d^4x_m d^4y_1 \cdots d^4y_n d^4u_1 \cdots d^4u_n d^4v_1 \cdots d^4v_n \\
&\times V_{\mu_1, \dots, \mu_n}^{(m,n)}(x_1, \dots, x_m, y_1, \dots, y_n, u_1 \cdots u_n, v_1 \cdots v_n) \\
&\times : \bar{\psi}(x_1) \psi(y_1) \cdots \bar{\psi}(x_m) \psi(y_m) \bar{\psi}(u_1) \gamma^{\mu_1} \psi(v_1) \cdots \bar{\psi} \gamma^{\mu_n} \psi(v_n) :, \quad (1)
\end{aligned}$$

where ψ is the quark field. The quantum transition amplitude is given by

$$Z_{\text{fi}} = \int D\psi D\bar{\psi} \exp \left(i \int d^4x L \right), \quad (2)$$

where L is the Lagrangian density of the system. Inserting the identity

$$1 = \int \prod_{x,y} D\Sigma_s(x,y) D\Sigma_\mu(x,y) D\sigma(x,y) D\omega^\mu(x,y) \exp\left(i \int dxdy \Sigma_s \{\sigma(x,y) - \bar{\psi}(x)\psi(y)\}\right) \times \exp\left(i \int dxdy \Sigma_\nu \{\bar{\psi}(x)\gamma^\nu \psi(y) - \omega^\nu(x,y)\}\right), \quad (3)$$

we introduce the auxiliary meson fields $\sigma(=\bar{\psi}\psi)$ and $\omega_\mu(=\bar{\psi}\gamma_\mu\psi)$ as well as the quark self-energies Σ_s and Σ_μ . In this model, the expectation values of σ and ω_0 fields are proportional to the quark scalar density and the baryon density, respectively.

Integrating the quark field, we obtain by means of the mean field approximation

$$Z_{\text{fi}} = \int D\sigma D\omega_\mu \exp(i\Gamma[\sigma, \omega_\mu]), \quad (4)$$

where Γ is the effective action and is given by

$$\Gamma[\sigma, \omega_\mu] = W_0[\Sigma_s[\sigma, \omega_\mu], \Sigma_\mu[\sigma, \omega_\mu]] - \sum_{m+n \geq 2} \int V_{\mu_1 \dots \mu_n}^{(m,n)} \sigma^m \omega^{\mu_1} \dots \omega^{\mu_n} + \text{Tr}(\sigma \Sigma_s[\sigma, \omega_\mu] - \omega_\mu \Sigma_\mu[\sigma, \omega_\mu]). \quad (5)$$

The W_0 represents the quark energy and the remaining parts represent the meson potential. The quark self-energies Σ_s and Σ_μ are determined by the following conditions.

$$\frac{\delta \Gamma}{\delta \sigma} = -\frac{\partial}{\partial \sigma} \left(\sum_{m+n \geq 2} \int V_{\mu_1 \dots \mu_n}^{(m,n)} \sigma^m \omega^{\mu_1} \dots \omega^{\mu_n} \right) + \Sigma_s = 0. \quad (6)$$

$$\frac{\delta \Gamma}{\delta \omega^\mu} = -\frac{\partial}{\partial \omega^\mu} \left(\sum_{m+n \geq 2} \int V_{\mu_1 \dots \mu_n}^{(m,n)} \sigma^m \omega^{\mu_1} \dots \omega^{\mu_n} \right) - \Sigma_\mu = 0. \quad (7)$$

3 Effective meson masses, effective couplings and EOS

Because of the conditions (6) and (7), the quark self-energies are strongly related to the meson potential. Therefore, at finite temperature and density, the effective meson-quark couplings are strongly related to the effective meson masses. In the uniform and rotationally invariant matter, we obtain

$$\frac{m_\sigma^{*2}}{m_\sigma^2} \equiv \frac{\partial^2 \epsilon}{\partial \sigma^2} = \frac{g_{s\sigma}^* \Pi g_{s\sigma}^*}{m_\sigma^2} + \frac{g_{s\sigma}^*}{g_\sigma} \quad \text{and} \quad \frac{m_\omega^{*2}}{m_\omega^2} \equiv -\frac{\partial^2 \epsilon}{\partial \omega_0^2} = -\frac{g_{s\omega}^* \Pi g_{s\omega}^*}{m_\omega^2} + \frac{g_{v\omega}^*}{g_\omega}, \quad (8)$$

where $g_{s\sigma}^* \equiv -\frac{\partial \Sigma_s}{\partial \sigma}$, $g_{s\omega}^* \equiv -\frac{\partial \Sigma_s}{\partial \omega_0}$, $g_{v\omega}^* \equiv -\frac{\partial \Sigma_0}{\partial \omega_0}$ and Π is the polarization function. The m_σ , m_ω , g_σ and the g_ω are the σ -meson mass, the ω -meson mass, the σ -quark coupling and the ω -quark coupling at zero temperature and zero density, respectively, and ϵ is the energy density of the system. If the effects of the mixing interaction, the term including $g_{s\omega}^*$, can be neglected, the square of the effective ω -meson mass is proportional to the effective ω -quark coupling. Therefore, the effective ω -quark coupling decreases as the effective ω -meson mass decreases.

In Fig. 1, we show the baryon density (ρ_B) dependence of the binding energy of nuclear matter at zero temperature. In the calculation, we assume that $g_{N\sigma}^* = 3g_{s\sigma}^*$ and $g_{N\omega}^* = 3g_{v\omega}^*$, where $g_{N\sigma}^*$ and $g_{N\omega}^*$ are the effective σ -nucleon and ω -nucleon couplings, respectively. In the nonlinear model (NLM) $g_{N\omega}^*/g_{N\omega} = m_\omega^{*2}/m_\omega^2 \sim 0.94$ at the normal density ρ_{B0} , whereas $g_{N\omega}^*/g_{N\omega} = m_\omega^{*2}/m_\omega^2 = 1$ in the linear model (LM). (See Fig. 2.) Although the effective ω -meson mass decreases in the NLM, the EOS in the NLM becomes much softer than that in the LM.

4 Summary

In summary, starting from the many quark interaction, we have constructed the nonlinear σ - ω model. The mesons are introduced as auxiliary fields. Effective quark-meson couplings are strongly related to effective meson masses, since they are derived simultaneously from the original many quark interactions. In this model, even if the effective ω -meson mass decreases due to the partial chiral restoration, the effective ω -quark (or nucleon) coupling decreases and the EOS of nuclear matter can become soft.

References

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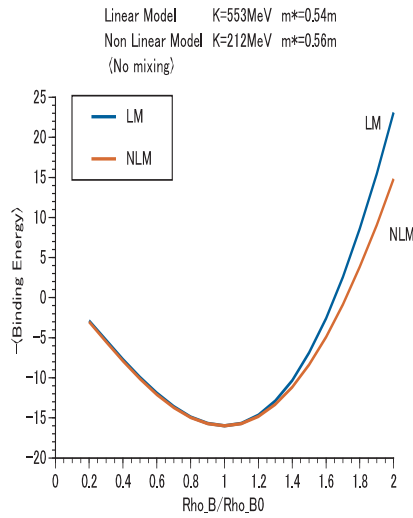


Fig.1 Binding energy (in MeV)
 K : Incompressibility
 m^* : Effective nucleon mass
at the normal density

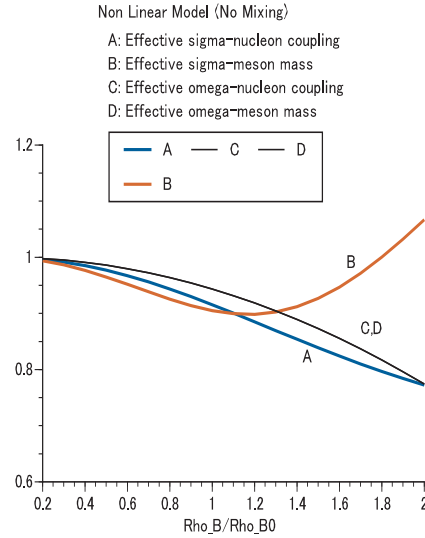


Fig.2 Squares of effective meson masses
and effective meson-nucleon
couplings (ratios) in NLM